LETTER TO THE EDITOR

Nearest-neighbour distribution function for systems of interacting particles

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Abstract. One of the basic quantities characterising a system of interacting particles is the nearest-neighbour distribution function H(r). We give a general expression for H(r) for a distribution of D-dimensional spheres which interact with an arbitrary potential. Specific results for H(r) are obtained, for the first time, for D-dimensional hard spheres with D=1, 2 and 3. Our results for D=3 are shown to be in excellent agreement with Monte Carlo computer-simulation data for a wide range of densities. From H(r), one can determine other quantities of fundamental interest such as the mean nearest-neighbour distance and the random close-packing density.

In considering systems composed of many interacting particles, a key fundamental question to ask is: what is the effect of the nearest neighbour on some reference particle in the system? The answer to this query requires knowledge of the nearest-neighbour distribution function H(r), i.e. the probability density associated with finding a nearest neighbour at some given distance r from the reference particle. From H(r) one can determine other quantities of fundamental interest such as the mean nearest-neighbour distance and the random close-packing density. Knowing H(r) is of importance in a host of problems in the physical and biological sciences, including liquids and amorphous solids [1-5], transport properties of suspensions and composite materials [6-8], stellar dynamics [9], and the structure of some cell membranes [10], to mention but a few examples. It should be emphasised that H(r) is different from the well known radial distribution function. The latter quantity is proportional to the probability of finding any particle (not necessarily the nearest one) a distance r away from a central particle.

Hertz [11] apparently was the first to consider the evaluation of H(r) for a system of 'point' particles, i.e. particles whose centres are randomly (Poisson) distributed. The *D*-dimensional generalisation of Hertz's [11] solution of H(r) for Poisson distributed points, at number density ρ , is given by

$$H(r) = \rho \frac{\mathrm{d}v_D(r)}{\mathrm{d}r} \exp[-\rho v_D(r)] \tag{1}$$

where $v_D(r)$ is the volume of a *D*-dimensional sphere of radius $r(v_1(r) = 2r, v_2(r) = \pi r^2, v_3(r) = \frac{4}{3}\pi r^3)$.

Interestingly, there is currently no theoretical formalism to obtain and compute H(r) for distributions of *finite-sized* interacting particles at arbitrary density[†]. In this letter, we briefly describe such general results for D-dimensional spheres. We then specifically determine H(r) and the mean nearest-neighbour distance for D-dimensional random arrays of impenetrable spheres of diameter σ as a function of density. (The rather lengthy derivation of all the theoretical results given here and the calculation of functions closely related to H(r) will be described in detail elsewhere [13].) The case D=1 (hard rods) may serve as a useful model of various types of layered media [14]. The case D=2 (hard discs) is a reasonable model of fibre-reinforced materials [15], thin films [15], certain types of cell membranes [10], etc. The case D=3 (hard spheres) has probably the widest application as it can be used to model liquids [1, 2, 16], amorphous solids [2-5], suspensions [6], porous media [7, 8], particulate composites [17], powders [18], etc.

We have derived an exact analytical representation of H(r) for homogeneous distributions of identical interacting D-dimensional spheres of diameter σ at number density ρ in terms of the so-called n-particle probability density functions $\rho_1, \rho_2, \ldots, \rho_n$. It is found [13] that

$$H(r) = \sum_{k=1}^{\infty} (-1)^{k+1} H^{(k)}(r)$$
 (2)

where

$$H^{(k)}(r) = \frac{1}{k!} \frac{\partial}{\partial r} \int \rho_{k+1}(\mathbf{R}^{k+1}) \prod_{i=2}^{k+1} m(|\mathbf{R}_1 - \mathbf{R}_i|; r) \, d\mathbf{R}_i$$
 (3)

with

$$m(y; r) = \begin{cases} 1 & y \le r \\ 0 & y > r, \end{cases} \tag{4}$$

The quantity $\rho_n(R_1, \ldots, R_n)$ characterises the probability of finding a configuration of n spheres with centres at positions $R^n = R_1, \ldots, R_n$, respectively, and is given information for the statistical ensemble under consideration. For spatially uncorrelated centres (Posson distribution), ρ_n is trivially a constant equal to ρ^n and our expression leads to the simple formula (1). On the other hand, if the particles are mutually impenetrable, then the ρ_n are generally quite complicated [16].

For the case of hard rods (D = 1), the ρ_n , for any n, are known exactly for equilibrium distributions [19]. Our relation for H then yields the exact dimensionless result

$$\sigma H(x) = \frac{2\eta}{1-\eta} \exp\left(\frac{-2\eta(x-1)}{1-\eta}\right) \qquad x > 1$$
 (5)

where $x = r/\sigma$ is a scaled distance and $\eta = \rho v_1(\sigma/2) = \rho \sigma$ is a reduced density. For x < 1, H(x) = 0 in any dimension.

For the cases of D=2 and D=3, however, the two-particle probability density ρ_2 (or equivalently, the radial distribution function) is only known approximately for

[†] The nearest-neighbour distribution function H(r) defined here should not be confused with the one defined by Reiss et al [12] in their scaled-particle theory. Whereas the former considers nearest neighbours around an actual inclusion centred at the origin, the latter considers nearest neighbours at a radial distance from the centre of a spherical cavity empty of sphere centres. The distinction between these two different types of nearest-neighbour distribution functions is fully detailed in [13].

arbitrary density, albeit accurately [16]; the higher-order $\rho_n(n \ge 3)$ are generally never known. This implies that an exact solution of H(r) for D=2 and 3 under general conditions is out of the question. For D=2 and 3, therefore, we have devised schemes to approximately sum the series using statistical mechanical theory [13] and found

$$\sigma H(x) = \frac{4\eta(2x-\eta)}{(1-\eta)^2} \exp\left(\frac{-4\eta}{(1-\eta)^2} [(x^2-1) + \eta(x-1)]\right) \qquad x > 1$$
 (6)

for hard discs (D=2), where $\eta = \rho v_2(\sigma/2)$, and

$$\sigma H(x) = 24\eta (ex^2 + fx + g) \exp\{-\eta [8e(x^3 - 1) + 12f(x^2 - 1) + 24g(x - 1)]\}$$
 $x > 1$ (7)

for hard spheres (D=3), where $\eta = \rho v_3(\sigma/2)$ and

$$e = \frac{1+\eta}{(1-\eta)^3} \qquad f = \frac{-\eta(3+\eta)}{2(1-\eta)^3} \qquad g = \frac{\eta^2}{2(1-\eta)^3}.$$
 (8)

It should be emphasised that the relations (5), (6), and (7) for D=1, D=2, and D=3, respectively, are new, i.e. it is the first time that expressions for H(r) valid for D-dimensional hard-sphere systems at arbitrary density have been given.

In figure 1 we plot H(r) for distributions of D-dimensional impenetrable spheres at a sphere volume fraction $\phi = \eta = 0.2$. Of course, for $r < \sigma$, H(r) = 0 for any D. For r near σ , the effect of increasing the dimensionality is to increase H(r), i.e. the likelihood of finding a nearest neighbour at such r increases with increasing D. Consistent with this behaviour is a decrease of H(r) with increasing D for large r.

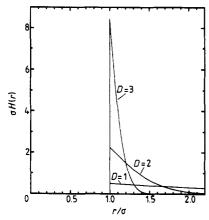


Figure 1. The dimensionless nearest-neighbour distribution function $\sigma H(r)$ for distributions of identical D-dimensional impenetrable spheres of diameter σ at a D-dimensional particle volume fraction $\phi=0.2$. Results for D=1,2 and 3 are obtained from (5), (6) and (7), respectively. For impenetrable spheres, the D-dimensional volume fraction ϕ equals the D-dimensional reduced density $\eta=\rho v_D(\sigma/2)$, where $v_D(r)$ is the D-dimensional volume of a sphere of radius r described in the text and ρ is the particle number density.

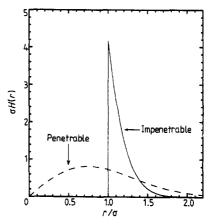


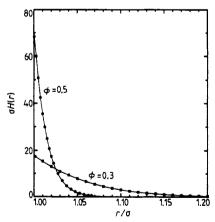
Figure 2. The dimensionless nearest-neighbour distribution function $\sigma H(r)$ for penetrable discs (Poisson distributed 'point' particles) and impenetrable discs of diameter σ as calculated from (1) and (6), respectively, at a particle area fraction $\phi=0.3$. For D-dimensional penetrable spheres, the sphere volume fraction $\phi=1-\exp(-\eta)$. Exclusion-volume effects associated with the hard cores considerably change the behaviour of h(r) relative to the idealised case of point particles. H(r) behaves qualitatively the same for these models in any dimension.

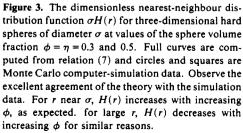
What is the effect of impenetrability of the spheres on H(r)? In figure 2 we compare Hertz's result (1) for Poisson distributed centres in two-dimensional space with our new result (6) for two-dimensional impenetrable discs at a disc area fraction $\phi = 0.2$. Note that exclusion-volume effects associated with hard cores lead to a nearest-neighbour distribution function which is strikingly different to the corresponding quantity for spatially uncorrelated discs. For $r < \sigma$, unlike hard discs, $H(r) \neq 0$ for penetrable discs since their centres can come arbitrarily close to one another. For large r, H(r) for penetrable discs is larger than H(r) for impenetrable discs since in the former system one is more likely to find larger 'void' regions surrounding the central particle as the result of interparticle overlap. The behaviour of H(r) for these models for any D is qualitatively the same.

Monte Carlo computer simulations in three dimensions have been carried out by Torquato and Lee [20] to obtain, among other quantities, H(r). A standard Metropolis [16] algorithm was employed to generate 200-6000 different realisations of 500 impenetrable spheres in a cubical cell with periodic boundary conditions. Figure 3 compares the simulation results with our relation (4) for $\phi = 0.2$ and $\phi = 0.5$. The agreement is seen to be excellent. In fact, one finds relatively good agreement up to $\phi = 0.6$, which is very close to the random close-packing volume fraction ϕ_c , estimated to range from 0.62-0.66 [2, 4]. In conclusion, this verifies the accuracy of the three-dimensional expression (7) (as well as the two-dimensional expression which is based on a similar approximation scheme) up to densities near the close-packing value (see discussion below).

Another important measure is the 'mean nearest-neighbour distance' I defined as

$$l = \int_{0}^{\infty} rH(r) \, \mathrm{d}r. \tag{9}$$





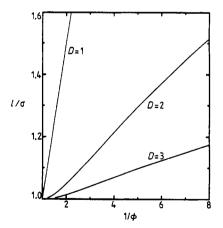


Figure 4. The dimensionless mean nearest-neighbour distance l/σ as a function of the inverse volume fraction ϕ^{-1} for distributions of *D*-dimensional impenetrable spheres with D=1, 2 and 3.

An operational definition for the random close-packing volume fraction ϕ_c , a quantity of great fundamental interest [2-5], then follows, i.e. the volume fraction at which $l = \sigma$. We have computed (9) for D-dimensional hard spheres using the exact formula (5) and the approximate relations (6) and (7) as a function of the D-dimensional inverse volume fraction ϕ^{-1} . These results are summarised in figure 4. As expected, at fixed ϕ , l increases with increasing D. Unlike our exact one-dimensional result which correctly predicts $\phi_c = 1$, our two-dimensional and three-dimensional results for l cannot correctly predict the 'critical' point ϕ_c . This is not surprising considering the difficulty of predicting ϕ_c for D=2 and 3 (heretofore this problem has defined an exact analytical solution) and because our approximations are 'mean field' in nature and hence cannot accurately predict critical points [5]. Our plots of l/σ as a function of ϕ^{-1} are approximately linear over the entire range of ϕ , except for the near vicinity of ϕ_c . Interestingly, extrapolation of these two-dimensional and three-dimensional data (using the linear range) to the limit $l/\sigma = 1$, yields values of ϕ_c which fall within the respective estimated ranges [4] (for D = 2, $\phi_c = 0.82 \pm 0.02$). Such linear extrapolations, however, are somewhat arbitrary. In future work we shall study methods for improving our approximations (6) and (7) in the near-critical region.

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