

# Bulk properties of composite media. I. Simplification of bounds on the shear modulus of suspensions of impenetrable spheres

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We study third-order upper and lower bounds on the shear modulus of a model composite made up of equisized, impenetrable spherical inclusions randomly distributed throughout a matrix phase. We determine greatly simplified expressions for the two key multidimensional cluster integrals (involving the three-point distribution function for one of the phases) arising in these bounds. These expressions are obtained by expanding the orientation-dependent terms in the integrand in spherical harmonics and employing the orthogonality property of this basis set. The resulting simplified integrals are in a form that makes them much easier to compute. The approach described here is quite general in the sense that it has application in cases where the spheres are permeable to one another (models of consolidated media such as sandstones and sintered materials) and to the determination of other bulk properties, such as the bulk modulus, thermal/electrical conductivity, and fluid permeability.

## I. INTRODUCTION

The problem we are generally concerned with is the theoretical prediction of the effective properties (transport, elastic, electromagnetic, etc.) of disordered composite media. This problem is of considerable fundamental and practical interest<sup>1-7</sup> and is exactly soluble, given the phase properties and the infinite set of  $n$ -point correlation functions<sup>8-10</sup> that statistically characterize the composite medium. The complete set of statistical functions is almost never known in practice, however. Under such circumstances one can either opt for some sort of approximate self-consistent scheme<sup>11-13</sup> or methods that enable us to place bounds on the effective property. Both of these methods have their own advantages and disadvantages and have been discussed elsewhere.<sup>5,6</sup>

We shall focus our attention on rigorous bounding techniques, since they provide a means of estimating the bulk property, given limited microstructural information on the heterogeneous material. Rigorous bounds are useful because (1) they enable one to test the merits of a theory; (2) one of the bounds can typically provide a relatively accurate estimate of the property<sup>7</sup>; and (3) as successively more microstructural information is included, the bounds become progressively tighter. The specific problem of interest in the present study is the determination of bounds (described below) on the effective shear modulus ( $G_e$ ) of a suspension of impenetrable equisized spherical inclusions.

Bounds on the effective elastic moduli that depend upon the  $n$ -point probability function  $S_n$  of the medium have been derived.<sup>14-17</sup> The  $S_n(\mathbf{x}^n)$  ( $\mathbf{x}^n \equiv \mathbf{x}_1, \mathbf{x}_2, \dots, \mathbf{x}_n$ ) give the probability of finding  $n$  points at positions  $\mathbf{x}^n$  all in one of the phases, for example, phase 2. For statistically homogeneous media  $S_1$  is simply equal to the volume fraction of phase 2,

$\phi_2$ . The second-order bounds on the elastic moduli of Hashin and Shtrikman<sup>14</sup> depend on  $S_1$  and, in a trivial way, on  $S_2$ . McCoy<sup>15</sup> obtained sharper third-order bounds (that were later simplified by Milton<sup>16</sup>), which involve information about  $S_1, S_2$ , and  $S_3$ . Subsequently, Milton and Phan-Thien<sup>17</sup> (MPT) derived third-order bounds, which improve upon the McCoy bounds, and new fourth-order bounds. Practical application of third- and fourth-order bounds on the effective elastic moduli has been very slow because of the difficulty involved in ascertaining  $S_3$  and  $S_4$  either theoretically or experimentally.

Employing the formalism of Torquato and Stell<sup>9,10</sup> to systematically represent and compute the  $S_n$ , third-order bounds on the effective elastic moduli have been recently computed for suspensions of fully penetrable spheres<sup>18</sup> and for dilute dispersions of spheres distributed with an arbitrary degree of penetrability.<sup>19</sup> More recently, third-order bounds on the effective bulk modulus of a composite with impenetrable spherical inclusions have been calculated.<sup>20</sup>

In this paper we consider the evaluation of the third-order McCoy and MPT bounds on the effective shear modulus  $G_e$  of a random distribution of impenetrable equisized spheres in a matrix. In Sec. II we present the MPT bounds and some relevant discussions. In Sec. III we invoke the diagrammatic notation<sup>9</sup> for the terms involved in  $S_3$  and present the two key integrals that operate on these terms. The complicated multidimensional cluster integrals are then simplified as far as possible by using expansions of the orientation-dependent quantities in spherical harmonics and the orthogonality of this basis set. The final result presented at the end of Sec. III is still nontrivial, but in a form that is now tractable on a computer. The computation, which is under progress, will be the subject of a subsequent paper. In passing

we may reiterate the fact that the method of attack presented here is quite general in the sense that it can be (and has been) readily and systematically applied to cases in which spheres are permeable to one another<sup>18</sup> and to the determination of other bulk properties, such as the conductivity and bulk modulus of composites,<sup>20</sup> and fluid permeability of porous media.<sup>21</sup> Finally, in Sec. IV we present our conclusions.

### III. BOUNDS

For any two-phase isotropic composite, McCoy<sup>15</sup> has derived bounds on the effective shear modulus  $G_e$ , given the shear moduli  $G_1, G_2$ , and the bulk moduli  $K_1$  and  $K_2$  of the respective phases, the volume fraction of one of the phases (say  $\phi_2 = 1 - \phi_1$ ), and several integrals involving derivatives of certain three-point correlation functions. Milton<sup>16</sup> later showed that the third-order McCoy bounds can be written in terms of  $\phi_2$  and only two multifold integrals  $\zeta_2$  and  $\eta_2$  (defined below) involving the three-point probability function  $S_3$ . Subsequently, MPT<sup>17</sup> derived improved third-order bounds on  $G_e$  that also depend upon  $\phi_2, \zeta_2$ , and  $\eta_2$ . For conciseness we present here only the MPT bounds, which read

$$\left( \langle G \rangle - \frac{6\phi_1\phi_2(G_1 - G_2)^2}{6\langle \bar{G} \rangle + \Xi^{-1}} \right) < G_e < \left( \langle G \rangle - \frac{6\phi_1\phi_2(G_1 - G_2)^2}{6\langle \bar{G} \rangle + \theta} \right), \quad (1)$$

where

$$\Xi = \frac{5\langle 1/G \rangle_\xi \langle 6/K - 1/G \rangle_\xi + \langle 1/G \rangle_\eta \langle 2/K + 21/G \rangle_\xi}{\langle 128/K + 99/G \rangle_\xi + 45\langle 1/G \rangle_\eta}, \quad (2)$$

$$\theta = \frac{3\langle G \rangle_\eta \langle 6K + 7G \rangle_\xi - 5\langle G \rangle_\xi^2}{\langle 2K - G \rangle_\xi + 5\langle G \rangle_\eta}, \quad (3)$$

and where the angular brackets denote the averages of the following types for any property  $b$ :

$$\begin{aligned} \langle b \rangle &= b_1\phi_1 + b_2\phi_2, \\ \langle b \rangle_\xi &= b_1\zeta_1 + b_2\zeta_2, \\ \langle b \rangle_\eta &= b_1\eta_1 + b_2\eta_2, \\ \langle \bar{b} \rangle &= b_1\phi_2 + b_2\phi_1. \end{aligned} \quad (4)$$

Here the quantities  $\zeta_i$  and  $\eta_i$  ( $i = 1, 2$ ) are the integrals (defined below) over the three-point probability function  $S_3$ . To be specific,  $S_n$  denotes the probability of finding  $n$  points in the particulate phase (here the impenetrable spheres). Then we have

$$\zeta_2 = 1 - \zeta_1 = (9/2\phi_1\phi_2)I[\hat{S}_3], \quad (5)$$

$$\eta_2 = 1 - \eta_1 = \frac{3}{2}\zeta_2 + (150/7\phi_1\phi_2)J[\hat{S}_3], \quad (6)$$

and

$$\hat{S}_3(r, s, t) = S_3(r, s, t) - S_2(r)S_2(s)/S_1, \quad (7)$$

where the integral operators  $I$  and  $J$  are defined as

$$I[f] = \frac{1}{8\pi^2} \int d\mathbf{r}_2 d\mathbf{r}_3 f(r_{12}, r_{13}, r_{23}) \frac{P_2(\hat{\mathbf{r}}_{12} \cdot \hat{\mathbf{r}}_{13})}{r_{12}^3 r_{13}^3}, \quad (8)$$

and

$$J[f] = \frac{1}{8\pi^2} \int d\mathbf{r}_2 d\mathbf{r}_3 f(r_{12}, r_{13}, r_{23}) \frac{P_4(\hat{\mathbf{r}}_{12} \cdot \hat{\mathbf{r}}_{13})}{r_{12}^3 r_{13}^3}, \quad (9)$$

$P_l$  being the Legendre polynomial of degree  $l$ . It may be mentioned that we operate the  $I$  and  $J$  on  $\hat{S}_3$  defined in (7) instead of  $S_3$  because it ensures the absolute convergence of these integrals.

It may further be noted that  $\zeta_i$  ( $i = 1$  or  $2$ ) is the only parameter needed for the third-order bounds on electrical conductivity or bulk modulus. The calculation of this parameter, and hence the evaluation of  $I[\hat{S}_3]$ , has already been considered in great detail in two previous papers.<sup>20</sup> The parameter  $\zeta_2$  is thus well tabulated for our needs, and hence we concern ourselves with the evaluation of the integral  $J[\hat{S}_3]$  or the parameter  $\eta_2$  only. The reader interested in the details of the  $\zeta_2$  calculation is referred to the above-mentioned papers.<sup>20</sup>

The evaluation of the bounds (1) becomes quite difficult partly because the three-point function  $S_3$  was not available theoretically or experimentally until recently and partly because of the complexity of the integral  $J$  [Eq. (9)]. Thus, to our knowledge, this is the first paper to deal with those bounds for the effective shear modulus for the case of a dispersion of equisized impenetrable spheres at arbitrary concentration in a matrix phase.

### III. SIMPLIFICATION OF $J[\hat{S}_3]$ FOR DISPERSIONS OF IMPENETRABLE SPHERES

Torquato and Stell<sup>9</sup> have shown that the  $n$ -point probability function for the *matrix* phase of dispersions of impenetrable spheres reduces to a finite series expansion in the density  $\rho$ , ending at the  $n$ th-order term. For our purposes we will use the notation of Ref. 9, except for the fact that  $S_n$  now will describe the  $n$ -point probability function for the *particle* phase, i.e., the quantity denoted by  $S_n$  in the Introduction.<sup>22</sup> As shown by Torquato and Stell,<sup>8</sup> given the  $n$ -point probability functions ( $S_1, S_2, \dots, S_n$ ) for one of the two phases, one can get any other  $n$ -point probability function and, in particular, the  $n$ -point probability function for the other phase. Thus, taking the expressions for the one-, two-, and three-point matrix probability functions,<sup>9</sup> we can write the three-point probability function in the particulate phase as

$$S_3 = S_3^{(1)}\phi_2 + S_3^{(2)}\phi_2^2 + S_3^{(3)}\phi_2^3, \quad (10)$$

where  $\phi_2 = \rho V_1$  is the particle volume fraction,  $V_1 = 4\pi a^3/3$  is the spherical volume of one inclusion,  $a$  is the radius of the sphere, and  $S_3^{(i)}$  stand for the following diagrams:

$$S_3^{(1)} = \frac{1}{V_1} \begin{array}{c} \bullet \\ \diagup \quad \diagdown \\ \circ \quad \circ \quad \circ \\ 1 \quad 2 \quad 3 \end{array}, \quad (11a)$$

$$S_3^{(2)} = \frac{1}{V_1^2} \left( \begin{array}{ccc} \bullet & \bullet & \bullet \\ \circ & \circ & \circ \\ 1 & 2 & 3 \end{array} + \begin{array}{ccc} \bullet & \bullet & \bullet \\ \circ & \circ & \circ \\ 1 & 3 & 2 \end{array} + \begin{array}{ccc} \bullet & \bullet & \bullet \\ \circ & \circ & \circ \\ 2 & 3 & 1 \end{array} \right) \quad (11b)$$

$$S_3^{(3)} = \frac{1}{V_1^3} \begin{array}{ccc} \bullet & \bullet & \bullet \\ \circ & \circ & \circ \\ 1 & 2 & 3 \end{array} \quad (11c)$$

Here the solid circles stand for dummy position vectors (of some spheres) that are to be integrated over the entire infinite volume,<sup>23</sup> the labeled open circles represent the position vectors  $r_1$ ,  $r_2$ , and  $r_3$  appearing in  $S_3$ , the broken line represents the bond

$$m(r) = \begin{cases} 1, & r < a, \\ 0, & r > a, \end{cases} \quad (12)$$

between the two positions involved, the solid line stands for the pair distribution function  $g_2 \equiv g$  of the spheres, and the crosshatched triangle for their triplet distribution function  $g_3$ .

For the calculation of shear modulus, we have to evaluate the functional (9) [evaluation of the other functional (8) has already been done<sup>20</sup>], which may be rewritten as

$$J[f] = \int_0^\infty \frac{dr_{12}}{r_{12}} \int_0^\infty \frac{dr_{13}}{r_{13}} \int_{-1}^1 d(\cos \theta_{213}) \times P_4(\cos \theta_{213}) f(r_{12}, r_{13}, r_{23}). \quad (13)$$

We have to evaluate the above for each of the diagrams in (11).  $P_4$  is the Legendre polynomial of degree 4, and  $\cos \theta_{jik} \equiv (\hat{r}_{ij} \cdot \hat{r}_{ik})$ .

### A. Evaluation of $J[S_3^{(1)}]$

The only diagram of (11a) may be evaluated by fixing the origin of coordinates at  $r_1$  and aligning the  $z$  axis along  $\hat{r}_{12}$ , as follows:

$$\begin{array}{ccc} \bullet \\ \circ & \circ & \circ \\ 1 & 2 & 3 \end{array} \equiv \int dr_4 m(r_{14}) m(r_{24}) m(r_{34}) = \int_0^\infty dr_{14} r_{14}^2 m(r_{14}) \int d\omega_{214} m(r_{24}) m(r_{34}), \quad (14)$$

where  $d\omega_{214} \equiv d(\cos \theta_{214}) d\phi$ . Following Barker and Monaghan,<sup>24</sup> we expand the angle-dependent functions in Legendre polynomials (more generally, in spherical harmonics). Thus, for example, for  $m(r_{24})$  we write

$$m(r_{24}) = m[(r_{12}^2 + r_{14}^2 - 2r_{12}r_{14} \cos \theta_{214})^{1/2}] = \sum_{l=0}^\infty M_l(r_{12}, r_{14}) P_l(\cos \theta_{214}), \quad (15)$$

where the expansion coefficients are given by (see Appendix A)

$$M_l(r_{12}, r_{14}) = \frac{2l+1}{2\pi^2} \int_0^\infty dk k^2 \tilde{m}(k) j_l(kr_{12}) j_l(kr_{14}). \quad (16)$$

$\tilde{m}(k)$  is the Fourier transform of  $m(r)$ , and  $j_l(x)$  is the spherical Bessel function of order  $l$ . For the Fourier trans-

form  $\tilde{f}(k)$  of a function  $f(r)$ , we will always use the definition

$$\tilde{f}(k) = \int dr f(r) \exp(ik \cdot r). \quad (17)$$

Similarly, we write  $m(r_{34})$  as

$$m(r_{34}) = \sum_{l,m} M_l(r_{13}, r_{14}) P_l(\cos \theta_{314}) = \sum_{l,m} \frac{4\pi}{2l+1} M_l(r_{13}, r_{14}) Y_{lm}^*(\omega_{213}) Y_{lm}(\omega_{214}), \quad (18)$$

using the addition theorem for spherical harmonics<sup>25</sup> in the second equality to bring out the specific angular variables needed. Employing the orthogonality of the spherical harmonics, we get

$$\int d\omega_{214} m(r_{24}) m(r_{34}) = \sum_l \frac{4\pi}{2l+1} M_l(r_{12}, r_{14}) M_l(r_{13}, r_{14}) P_l(\cos \theta_{213}), \quad (19)$$

and hence

$$J \left( \begin{array}{ccc} \bullet \\ \circ & \circ & \circ \\ 1 & 2 & 3 \end{array} \right) = \frac{8\pi}{81} \int_0^\infty dr r^2 m(r) \left( \int_0^\infty \frac{ds}{s} M_4(s, r) \right)^2. \quad (20)$$

To calculate the second integral above, we may use (16) and the fact that

$$\tilde{m}(k) = (4\pi a^2/k) j_1(ka) \quad (21)$$

to obtain  $M_4$ . Thus, we find

$$\begin{aligned} \int_0^\infty \frac{ds}{s} M_4(s, r) &= \frac{18a^2}{\pi} \int_0^\infty dk k j_1(ka) j_4(kr) \int_0^\infty \frac{ds}{s} j_4(ks) \\ &= \frac{12a^2}{5\pi} \int_0^\infty dk k j_1(ka) j_4(kr) \\ &= \left[ 3 \left( \frac{a}{r} \right)^3 - \frac{21}{5} \left( \frac{a}{r} \right)^5 \right] H(r-a), \end{aligned} \quad (22)$$

where  $H(x)$  is the Heaviside unit function

$$H(x) = \begin{cases} 0, & x < 0, \\ 1, & x > 0, \end{cases} \quad (23)$$

and the integrals involving the  $j_l$ 's may be found in Ref. 26. Now we can see that the contribution of the diagram in (11a) to  $J$  is zero because (20) has conflicting step-function requirements in its two integrals. Thus, we find that

$$J[S_3^{(1)}] = 0. \quad (24)$$

### B. Evaluation of $J[S_3^{(2)}]$

To simplify the contribution of the diagrams in (11b), we utilize the freedom afforded by the homogeneity and isotropy of the system to conveniently choose the origin and orientation of the coordinate frame. If we first choose the origin at  $r_4$ , then we find for the first term of  $S_3^{(2)}$  in (11b) that

$$J\left(\frac{1}{V_1^2} \begin{array}{c} \bullet \text{---} \bullet \\ \diagdown \quad \diagup \\ \circ_1 \quad \circ_2 \quad \circ_3 \end{array}\right) = \frac{1}{8\pi^2 V_1^2} \int d\mathbf{r}_1 d\mathbf{r}_2 d\mathbf{r}_3 \frac{P_4(\cos \theta_{213})}{r_{12}^3 r_{13}^3} \int d\mathbf{r}_5 m(r_{14})m(r_{24})g(r_{45})m(r_{35})$$

$$= \frac{1}{8\pi^2} \int d\mathbf{r}_5 g(r_{45})W_1(r_{45}), \quad (25)$$

where

$$W_1(r_{45}) = \frac{1}{V_1^2} \int d\mathbf{r}_1 d\mathbf{r}_2 d\mathbf{r}_3 \frac{P_4(\cos \theta_{213})}{r_{12}^3 r_{13}^3} m(r_{14})m(r_{24})m(r_{35}). \quad (26)$$

Using (18) for the expansion of  $m(r_{24})$  and  $m(r_{35})$ , and then using the completeness relation for the  $Y_{lm}$ 's, we get

$$W_1(r_{45}) = \frac{1}{V_1^2} \int d\mathbf{r}_1 d\mathbf{r}_2 m(r_{14})m(r_{24}) \frac{1}{r_{12}^3} \sum_{l,m} \frac{4\pi}{2l+1} \int_0^\infty \frac{dr_{13}}{r_{13}} M_l(r_{13}, r_{15}) \int d\omega_{213} P_4(\cos \theta_{213}) Y_{lm}^*(\omega_{213}) Y_{lm}(\omega_{215})$$

$$= \frac{1}{V_1^2} \int d\mathbf{r}_1 m(r_{14}) \left(\frac{4\pi}{9} \int_0^\infty \frac{dx}{x} M_4(x, r_{15})\right) \int d\mathbf{r}_2 \frac{P_4(\cos \theta_{215})}{r_{12}^3} m(r_{24})$$

$$= \frac{1}{V_1^2} \int d\mathbf{r}_1 m(r_{14}) P_4(\cos \theta_{415}) \left(\frac{4\pi}{9} \int_0^\infty \frac{dx}{x} M_4(x, r_{15})\right) \left(\frac{4\pi}{9} \int_0^\infty \frac{dy}{y} M_4(y, r_{14})\right)$$

$$= \frac{1}{V_1^2} \int d\mathbf{r}_1 H(a - r_{14}) P_4(\cos \theta_{415}) \left\{ \frac{4}{3} \pi \left[ \left(\frac{a}{r_{15}}\right)^3 - \frac{7}{5} \left(\frac{a}{r_{15}}\right)^5 \right] H(r_{15} - a) \right\}$$

$$\times \left\{ \frac{4}{3} \pi \left[ \left(\frac{a}{r_{14}}\right)^3 - \frac{7}{5} \left(\frac{a}{r_{14}}\right)^5 \right] H(r_{14} - a) \right\}$$

$$= 0, \quad (27)$$

because of the conflicting demands of the step functions on  $r_{14}$ . In deriving (27), we have made use of (22). Thus

$$J\left(\frac{1}{V_1^2} \begin{array}{c} \bullet \text{---} \bullet \\ \diagdown \quad \diagup \\ \circ_1 \quad \circ_2 \quad \circ_3 \end{array}\right) = 0. \quad (28)$$

Now interchanging labels 2 and 3 in the procedure above gives the same integral, and hence the contribution of the second diagram of (11b) to  $J[S_3^{(2)}]$  is also zero. But

such is not the case for the last diagram of (11b). Proceeding as above, one obtains

$$J\left(\frac{1}{V_1^2} \begin{array}{c} \bullet \text{---} \bullet \\ \diagdown \quad \diagup \\ \circ_1 \quad \circ_2 \quad \circ_3 \end{array}\right) = \frac{1}{8\pi^2} \int d\mathbf{r}_5 g(r_{45})W_2(r_{45}), \quad (29)$$

where

$$W_2(r_{45}) = \frac{1}{V_1^2} \int d\mathbf{r}_1 d\mathbf{r}_2 d\mathbf{r}_3 \frac{P_4(\cos \theta_{213})}{r_{12}^3 r_{13}^3} m(r_{15})m(r_{24})m(r_{34})$$

$$= \frac{4\pi}{a^6} \int_0^\infty dr_{14} r_{14}^2 M_0(r_{14}, r_{45}) \left[ \left(\frac{a}{r_{14}}\right)^3 - \frac{7}{5} \left(\frac{a}{r_{14}}\right)^5 \right]^2 H(r_{14} - a). \quad (30)$$

In obtaining (30) above we have made repeated use of (22). Next, using the inverse expansion of (15) to write  $M_0$  and making a change of variables back to  $r_{15}$ , we get

$$M_0(r_{14}, r_{45}) = \frac{1}{2} \int_{-1}^1 d(\cos \theta)$$

$$\times m[(r_{14}^2 + r_{45}^2 - 2r_{14}r_{45} \cos \theta)^{1/2}]$$

$$= \frac{1}{2} \int_{|r_{14}-r_{45}|}^{r_{14}+r_{45}} dr_{15} \frac{r_{15}}{r_{14} r_{45}} m(r_{15}). \quad (31)$$

Use of (31) in (30) leads to

$$W_2(t) = A - B + C, \quad (32)$$

where

$$\begin{Bmatrix} A \\ B \\ C \end{Bmatrix} = \begin{Bmatrix} 1 \\ \frac{14}{5} a^2 \\ \frac{49}{25} a^4 \end{Bmatrix} \frac{2\pi}{t} \int_a^\infty dr \begin{Bmatrix} \frac{1}{r^5} \\ \frac{1}{r^7} \\ \frac{1}{r^9} \end{Bmatrix} \int_{|r-t|}^{r+t} ds sm(s). \quad (33)$$

Now, if we perform the two integrals in (33), we finally obtain

$$J[S_3^{(2)}] = \frac{1}{2\pi} \int_\sigma^\infty dr r^2 g(r) W_2(r), \quad (34)$$

where for  $r > \sigma$

$$W_2(r) = \frac{4}{3} \pi \left( \frac{a^3}{(r^2 - a^2)^3} - \frac{14}{5} \frac{a^5}{(r^2 - a^2)^4} - \frac{63}{25} \frac{a^7}{(r^2 - a^2)^5} + \frac{196}{25} \frac{a^9}{(r^2 - a^2)^6} + \frac{168}{25} \frac{a^{11}}{(r^2 - a^2)^7} \right). \quad (35)$$

Here  $\sigma = 2a$  is the sphere diameter and we have used the fact that  $g(r) = 0$  for  $r < \sigma$ . It may be noted that in the low-density limit  $\rho \rightarrow 0$ ,  $g(r) \rightarrow 1$  for  $r > \sigma$ , and then (34) reduces to

$$\lim_{\rho \rightarrow 0} J[S_3^{(2)}] = \frac{7213}{109350} - \frac{1}{24} \ln 3. \quad (36)$$

The importance of this result lies in the fact that if we write a low-density expansion of the parameter  $\eta_2$  as

$$\eta_2 = f_1 \phi_2 + O(\phi_2^2); \quad (37)$$

one can then show that

$$f_1 = \lim_{\rho \rightarrow 0} \frac{45}{42} J[S_3^{(2)}] + \lim_{\rho \rightarrow 0} \frac{150}{7} J[S_3^{(2)}]. \quad (38)$$

In a previous paper<sup>27</sup> it has been shown that

$$\lim_{\rho \rightarrow 0} I[S_3^{(2)}] = \frac{5}{24} - \frac{1}{24} \ln 3. \quad (39)$$

$$\int dr_2 m(r_{25}) \int dr_3 m(r_{36}) \frac{P_4(\cos \theta_{213})}{r_{12}^3 r_{13}^3} = V_1^2 H(r_{15} - a) H(r_{16} - a) \left[ 1 - \frac{7}{5} \left( \frac{a}{r_{15}} \right)^2 \right] \left[ 1 - \frac{7}{5} \left( \frac{a}{r_{16}} \right)^2 \right] \frac{P_4(\cos \theta_{516})}{r_{15}^3 r_{16}^3}. \quad (42)$$

In Eqs. (40) and (41),  $r_{14} < a$  and  $r_{46} > 2a$  because of the  $m$  and  $g$  functions involved. Thus  $H(r_{16} - a)$  is redundant in (42). Similarly, because  $r_{14} < a$  and  $r_{45} > 2a$ , the use of  $H(r_{15} - a)$  is redundant as well. Thus, we drop these  $H$  functions, and Eq. (41) is simplified to

$$Q(r_{45}, r_{46}, r_{56}) = \frac{2}{V_1} \int dr_1 m(r_{14}) \left[ 1 - \frac{7}{5} \left( \frac{a}{r_{15}} \right)^2 \right] \times \left[ 1 - \frac{7}{5} \left( \frac{a}{r_{16}} \right)^2 \right] \frac{P_4(\cos \theta_{516})}{r_{15}^3 r_{16}^3}. \quad (43)$$

The difficulty in simplifying this expression any further lies in explicitly bringing out the orientation dependence of the integrand for the final integration over  $r_1$ . For this we shall use the coordinate frame arrangement shown in Fig. 1. With

$$Q(r_{45}, r_{46}, r_{56}) = \frac{2}{V_1} \sum_{m=0}^4 \alpha_m \frac{(4-m)!}{(4+m)!} \int_0^a dr_{14} r_{14}^2 \int d\omega_{641} \left\{ \frac{P_4^m(\cos \theta_{415})}{r_{15}^3} \left[ 1 - \frac{7}{5} \left( \frac{a}{r_{15}} \right)^2 \right] \right\} \times \left\{ \frac{P_4^m(\cos \theta_{416})}{r_{16}^3} \left[ 1 - \frac{7}{5} \left( \frac{a}{r_{16}} \right)^2 \right] \right\} \cos(m\psi). \quad (47)$$

We show in Appendix B that each of the expressions within large brackets in (47) can be expanded in terms of the corresponding opposite angles at the base of the coordinate frame, giving

$$Q(r_{45}, r_{46}, r_{56}) = \frac{2}{V_1} \sum_{m=0}^4 \alpha_m \frac{(4-m)!}{(4+m)!} \int_0^a dr_{14} r_{14}^2 \int d\omega_{641} \times \sum_{l,l'} \left( \frac{r_{14}^{l-4}}{r_{45}^{l-1}} \beta_{lm} - \frac{r_{14}^{l-2}}{r_{45}^{l+1}} \gamma_{lm} - \frac{7}{5} a^2 \frac{r_{14}^{l-4}}{r_{45}^{l+1}} \nu_{lm} \right) P_l^m(\cos \theta_{541}) \times \left( \frac{r_{14}^{l'-4}}{r_{46}^{l'-1}} \beta_{l'm} - \frac{r_{14}^{l'-2}}{r_{46}^{l'+1}} \gamma_{l'm} - \frac{7}{5} a^2 \frac{r_{14}^{l'-4}}{r_{46}^{l'+1}} \nu_{l'm} \right) P_{l'}^m(\cos \theta_{641}) \cos(m\psi), \quad (48)$$

Combining (36) and (39) in (38), one finds that  $f_1 = 0.48274$ , which is the result quoted in the previously mentioned paper<sup>19</sup> on the third-order bounds on shear modulus in the dilute limit.

### C. Evaluation of $J[S_3^{(3)}]$

To simplify the contribution of the diagram in (11c), we employ the same technique used in the previous subsection. Thus, choosing the origin at  $r_4$ , we find that

$$J[S_3^{(3)}] = \frac{1}{16\pi^2} \int dr_5 dr_6 g_3(r_{45}, r_{46}, r_{56}) \times Q(r_{45}, r_{46}, r_{56}), \quad (40)$$

where

$$Q(r_{45}, r_{46}, r_{56}) = \frac{2}{V_1^3} \int dr_1 dr_2 dr_3 m(r_{14}) m(r_{25}) \times m(r_{36}) \frac{P_4(\cos \theta_{213})}{r_{12}^3 r_{13}^3}. \quad (41)$$

The integrals over  $r_3$  first and then  $r_2$  are done by following the same method as applied to simplify (26) to the form (27). The result in this case is

respect to this figure, we use the identity

$$\cos \theta_{516} = \cos \theta_{415} \cos \theta_{416} + \sin \theta_{415} \sin \theta_{416} \cos \psi, \quad (44)$$

where  $\psi$  is the angle between the planes 541 and 641, to write the addition theorem expansion for  $P_4(\cos \theta_{516})$  as

$$P_4(\cos \theta_{516}) = \sum_{m=0}^4 \alpha_m \frac{(4-m)!}{(4+m)!} P_4^m(\cos \theta_{415}) \times P_4^m(\cos \theta_{416}) \cos(m\psi), \quad (45)$$

where

$$\alpha_m = 1, \quad m = 0, \\ = 2, \quad m > 0. \quad (46)$$

Using (45) in (43), we find that

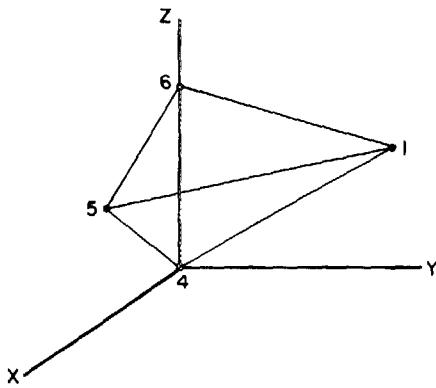


FIG. 1. Coordinate system for Eq. (43).

where

$$\begin{aligned} \beta_{10} &= \frac{7(l-3)(l-2)(l-1)l}{24(2l-1)}, \\ \beta_{11} &= -\frac{7(l-3)(l-2)(l-1)}{6(2l-1)}, \\ \beta_{12} &= \frac{7(l-3)(l-2)}{2(2l-1)}, \\ \beta_{13} &= -\frac{7(l-3)}{2l-1}, \\ \beta_{14} &= \frac{7}{2l-1}, \end{aligned} \quad (49)$$

$$\begin{aligned} \gamma_{10} &= \frac{(l-2)(l-1)l(7l+11)}{24(2l+3)}, \\ \gamma_{11} &= -\frac{(l-2)(l-1)(7l+9)}{6(2l+3)}, \\ \gamma_{12} &= \frac{(l-2)(7l+3)}{2(2l+3)}, \\ \gamma_{13} &= -\frac{7(l-1)}{2l+3}, \\ \gamma_{14} &= \frac{7}{2l+3}, \end{aligned} \quad (50)$$

and

$$v_{1m} = \left(\frac{2l-1}{7}\right)\beta_{1m}. \quad (51)$$

To be able to do the angular integral in (48), we rotate the coordinate frame of Fig. 1 by an angle  $\theta_{641}$  about an axis perpendicular to the  $(\hat{r}_{46}, \hat{r}_{41})$  plane with a temporary reassignment of the coordinate frame so that  $\hat{r}_{41}$  is in the  $(x, z)$  plane as in Fig. 2. The unit vector  $\hat{r}_{45}$  has an orientation  $(\theta_{546}, \phi)$  with respect to the  $(x, y, z)$  frame and  $(\theta_{541}, \psi)$  with respect to the rotated  $(x', y', z')$  frame, because  $\psi$  is the angle between the  $(x, z)$  plane and the  $(\hat{r}_{41}, \hat{r}_{51})$  plane. The Euler angles of rotation between the frames are  $(0, \theta_{641}, 0)$ . Thus the transformation theorem for this special case is<sup>25</sup>

$$Y_{lm}(\theta_{541}, \psi) = \sum_{m'} d_{m', m}^l(\theta_{641}) Y_{lm'}(\theta_{546}, \phi). \quad (52)$$

Writing out the spherical harmonics on both the sides in terms of the associated Legendre functions, we get

$$P_l^m(\cos \theta_{541}) \cos(m\psi) = \left(\frac{(l+m)!}{(l-m)!}\right)^{1/2} \sum_{m'=0}^l \alpha_{m'} (-1)^{m'-m} \left(\frac{(l-m')!}{(l+m')!}\right)^{1/2} d_{m', m}^l(\theta_{641}) P_l^{m'}(\cos \theta_{546}) \cos(m'\phi). \quad (53)$$

Using this, we find the angular integral in (48) to be

$$\int d(\cos \theta_{641}) d\phi [P_l^m(\cos \theta_{541}) \cos(m\psi)] P_l^m(\cos \theta_{641}) = \frac{4\pi}{2l+1} \frac{(l+m)!}{(l-m)!} P_l(\cos \theta_{546}) \delta_{l, l}, \quad (54)$$

and thus (48) reduces to

$$\begin{aligned} Q(r_{45}, r_{46}, r_{56}) &= \frac{2}{V_1} \sum_{m=0}^4 \alpha_m \frac{(4-m)!}{(4+m)!} \int_0^\infty dr_{14} r_{14}^2 \sum_l \frac{4\pi}{2l+1} \frac{(l+m)!}{(l-m)!} \left( \frac{r_{14}^{l-4}}{r_{45}^{l-1}} \beta_{lm} - \frac{r_{14}^{l-2}}{r_{45}^{l+1}} \gamma_{lm} - \frac{7}{5} a^2 \frac{r_{14}^{l-4}}{r_{45}^{l+1}} v_{lm} \right) \\ &\quad \times \left( \frac{r_{14}^{l-4}}{r_{46}^{l-1}} \beta_{lm} - \frac{r_{14}^{l-2}}{r_{46}^{l+1}} \gamma_{lm} - \frac{7}{5} a^2 \frac{r_{14}^{l-4}}{r_{46}^{l+1}} v_{lm} \right) P_l(\cos \theta_{546}). \end{aligned} \quad (55)$$

Next we compute the integral over  $r_{14}$  to obtain

$$\begin{aligned} Q(r_{45}, r_{46}, r_{56}) &= \frac{2}{V_1} \sum_{m=0}^4 \alpha_m \frac{(4-m)!}{(4+m)!} \sum_l \frac{4\pi}{2l+1} \frac{(l+m)!}{(l-m)!} \left[ \frac{a^{2l-1}}{r_{45}^{l+1} r_{46}^{l+1}} \left( \frac{\gamma_{lm}^2}{2l-1} + \frac{14}{5} \frac{\gamma_{lm} v_{lm}}{2l-3} + \frac{49}{25} \frac{v_{lm}^2}{2l-5} \right) \right. \\ &\quad \left. - a^{2l-3} \left( \frac{1}{r_{45}^{l+1} r_{46}^{l-1}} + \frac{1}{r_{45}^{l-1} r_{46}^{l+1}} \right) \left( \frac{\beta_{lm} \gamma_{lm}}{2l-3} + \frac{7}{5} \frac{\beta_{lm} v_{lm}}{2l-5} \right) + \frac{a^{2l-5}}{r_{45}^{l-1} r_{46}^{l-1}} \frac{\beta_{lm}^2}{2l-5} \right] P_l(\cos \theta_{546}). \end{aligned} \quad (56)$$

In the above we complete the sum over  $m$ , using (46) and (49)–(51), and then after a considerable amount of algebra, find that

$$\begin{aligned} Q(r_{45}, r_{46}, r_{56}) &= \frac{14}{5!} \sum_{l=4}^\infty l(l-1)(l-2)(l-3) \frac{(2l-3)}{(2l-1)} \frac{a^{2l-8}}{r_{45}^{l-1} r_{46}^{l-1}} \\ &\quad \times \left[ 1 - \frac{2}{5} (l+1) \left( \frac{2l-1}{2l-3} \right) \left( \frac{a}{r_{45}} \right)^2 \right] \left[ 1 - \frac{2}{5} (l+1) \left( \frac{2l-1}{2l-3} \right) \left( \frac{a}{r_{46}} \right)^2 \right] P_l(\cos \theta_{546}) \\ &\quad + \frac{8}{5!} \sum_{l=3}^\infty \frac{l(l-1)(l-2)(11l+15)}{(2l+3)(2l-3)} \frac{a^{2l-4}}{r_{45}^{l+1} r_{46}^{l+1}} P_l(\cos \theta_{546}). \end{aligned} \quad (57)$$

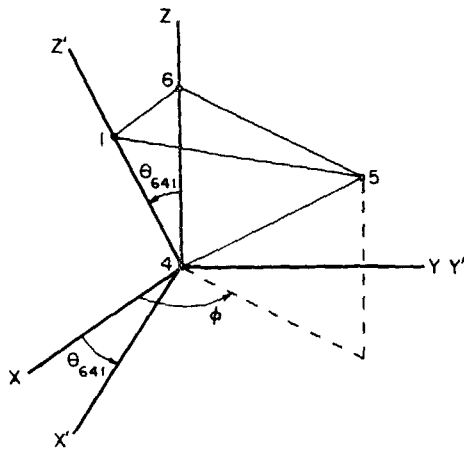


FIG. 2. Coordinate system for Eq. (52).

The result (57), when substituted in (40), provides the final simplified form of  $J[S_3^{(3)}]$ . Since the only orientation dependence in  $Q(r_{45}, r_{46}, r_{56})$  comes through  $P_l(\cos \theta_{546})$ , that is, through the angle between  $r_{45}$  and  $r_{46}$ , it is clear that if we replace  $g_3$  in (40) by  $g(r_{45})g(r_{46})$ , then the integral

$$I[\hat{S}_3] = \frac{2}{3} \phi_2^2 a^3 \int_{2a}^{\infty} dr \frac{r^2 g(r)}{(r^2 - a^2)^3} + \frac{\phi_2^3}{16\pi^2} \sum_{l=2}^{\infty} l(l-1)a^{2l-4} \int d\mathbf{r}_2 d\mathbf{r}_3 [g_3(r_{12}, r_{13}, r_{23}) - g(r_{12})g(r_{13})] \frac{P_l(\cos \theta_{213})}{r_{12}^{l+1} r_{13}^{l+1}}. \quad (59)$$

#### IV. CONCLUSIONS

For the model of impenetrable equisized spherical inclusions randomly distributed throughout a matrix, we have now simplified expressions for the two key integrals  $I[\hat{S}_3]$  and  $J[\hat{S}_3]$  that arise in the third-order McCoy and MPT bounds on the effective shear modulus  $G_e$ . It may again be noted that the simplification of  $I[\hat{S}_3]$  was done in a previous paper.<sup>20</sup> Both of these tasks were accomplished by expanding orientation-dependent terms in the two integrands in spherical harmonics and utilizing the orthogonality properties of this basis set. The resulting simplified integrals are shown to depend upon the one-, two-, and three-body distribution functions. We believe that this is the first time that the key integral  $J[\hat{S}_3]$ , required for the third-order bounds on  $G_e$ , has been simplified to this extent. In a subsequent paper we shall employ this simplified form for  $J$  to compute the McCoy and MPT bounds for a wide range of sphere volume fractions.

#### ACKNOWLEDGMENTS

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would vanish identically. But, whereas the integral (40) is conditionally convergent, depending upon the coordinate system and the order of performing the integration, the subtraction of the  $g(r_{45})g(r_{46})$  term from  $g_3$  before doing the integral makes it absolutely convergent. This fact was already remarked upon following Eq. (9).

If we combine the results from the previous three subsections, namely, (24) and (34) along with (35), and (40) along with (57), then the key integral  $J[\hat{S}_3]$  takes the form

$$J[\hat{S}_3] = \frac{1}{8\pi^2} \phi_2^2 \int d\mathbf{r} g(r) W_2(r) + \frac{\phi_2^3}{16\pi^2} \int d\mathbf{r}_2 d\mathbf{r}_3 [g_3(r_{12}, r_{13}, r_{23}) - g(r_{12})g(r_{13})] Q(r_{12}, r_{13}, r_{23}), \quad (58)$$

where the function  $Q$  is given in (57) and where  $W_2(r)$  is given in (35). Finally, for completeness, we end this section by giving the simplified form of the other key integral  $I[\hat{S}_3]$  as obtained by Lado and Torquato<sup>20</sup>:

#### APPENDIX A

As already used in (15), and following Barker and Monaghan,<sup>24</sup> we expand angle-dependent functions  $f(r_{23})$ , which are well behaved (i.e., functions with a finite number of finite discontinuities), in Legendre polynomials:

$$f(r_{23}) = \sum_{l=0}^{\infty} F_l(r_{12}, r_{13}) P_l(\cos \theta_{213}), \quad (A1)$$

where the orthogonality of the Legendre polynomials lead to the inverse expansion

$$F_l(r_{12}, r_{13}) = \frac{2l+1}{2} \int_{-1}^1 d(\cos \theta_{213}) f(r_{23}) P_l(\cos \theta_{213}), \quad (A2)$$

and

$$r_{23}^2 = r_{12}^2 + r_{13}^2 - 2r_{12}r_{13} \cos \theta_{213}. \quad (A3)$$

But  $f(r_{23})$  may also be expanded in plane waves as

$$f(r_{23}) = \frac{1}{(2\pi)^3} \int d\mathbf{k} \tilde{f}(k) \exp(i\mathbf{k} \cdot \mathbf{r}_{23}), \quad (A4)$$

where the Fourier transform  $\tilde{f}(k)$  is given by an expression similar to (17). If we now arrange that  $r_1$  is the origin,  $\hat{r}_{12}$  is along the  $z$  axis, the  $(\hat{r}_{12}, \hat{r}_{13})$  plane is the  $(x, z)$  plane, and let  $(\theta, \phi)$  be the angular coordinates of the wave vector  $\mathbf{k}$  in this frame, then the well-known expansion of plane waves in spherical waves gives<sup>28</sup>

$$\exp(i\mathbf{k} \cdot \mathbf{r}_{23}) = \exp(i\mathbf{k} \cdot \mathbf{r}_{13}) \exp(-ikr_{12} \cos \theta)$$

$$= \left( 4\pi \sum_{l,m} i^l j_l(kr_{13}) Y_{lm}^*(\theta_{213}, 0) Y_{lm}(\theta, \phi) \right) \left( \sum_{l'} (2l'+1) (-i)^{l'} j_{l'}(kr_{12}) P_{l'}(\cos \theta) \right). \quad (A5)$$

Then, using (A5) in the plane-wave expansion of  $f(r_{23})$ , one gets

$$f(r_{23}) = \frac{1}{2\pi^2} \sum_{l=0}^{\infty} (2l+1) P_l(\cos \theta_{213}) \times \int_0^{\infty} dk k^2 \tilde{f}(k) j_l(kr_{12}) j_l(kr_{13}). \quad (\text{A6})$$

Comparison of (A6) with (A1) gives

$$F_l(r_{12}, r_{13}) = \frac{2l+1}{2\pi^2} \int_0^{\infty} dk k^2 \tilde{f}(k) j_l(kr_{12}) j_l(kr_{13}), \quad (\text{A7})$$

a result we made use of in (20).

## APPENDIX B

This appendix deals with the specific task of rewriting  $P_4^m(\cos \phi)/t^3$  and  $P_4^m(\cos \phi)/t^5$  (see Fig. 3) in terms of  $r, s$ , and  $\cos \theta$  for  $m = 0, 1, 2, 3$ , and 4, and the condition  $s < r$ . We start with the generating function of the Legendre polynomials,<sup>29</sup>

$$\frac{r}{t} = \left(1 - 2 \frac{s}{r} \cos \theta + \frac{s^2}{r^2}\right)^{-1/2} = \sum_{l=0}^{\infty} \left(\frac{s}{r}\right)^l P_l(\cos \theta). \quad (\text{B1a})$$

Successive differentiation of (B1a) with respect to  $\cos \theta$ , denoted by primes, gives

$$\left(\frac{r}{t}\right)^3 = \sum_l \left(\frac{s}{r}\right)^{l-1} P_l'(\cos \theta), \quad (\text{B1b})$$

$$\left(\frac{r}{t}\right)^5 = \frac{1}{3} \sum_l \left(\frac{s}{r}\right)^{l-2} P_l''(\cos \theta), \quad (\text{B1c})$$

$$\left(\frac{r}{t}\right)^7 = \frac{1}{15} \sum_l \left(\frac{s}{r}\right)^{l-3} P_l'''(\cos \theta), \quad (\text{B1d})$$

(i)  $m = 0$  case

For the  $m = 0$  case we have

$$\begin{aligned} \left(\frac{r}{t}\right)^3 P_4(\cos \phi) &= \frac{35}{8} \left(\frac{r}{t}\right)^7 \sin^4 \theta - 5 \left(\frac{r}{t}\right)^5 \sin^2 \theta + \left(\frac{r}{t}\right)^3 \\ &= \sum_l \left(\frac{s}{r}\right)^{l-3} \left(\frac{7}{24} (1-x^2)^2 P_l''' - \frac{5}{3} (1-x^2) P_l''_{l-1} + P_l'_{l-2}\right) \\ &= \sum_l \left(\frac{s}{r}\right)^{l-3} \frac{1}{24} \frac{(l-1)(l-2)}{(2l+1)} [7l(l+1)P_{l+1} - (l-3)(7l+4)P_{l-1}] \\ &= \sum_l \left[\left(\frac{s}{r}\right)^{l-4} \frac{7l(l-1)(l-2)(l-3)}{24(2l-1)} - \left(\frac{s}{r}\right)^{l-2} \frac{l(l-1)(l-2)(7l+11)}{24(2l+3)}\right] P_l \end{aligned} \quad (\text{B4})$$

and

$$\begin{aligned} \left(\frac{r}{t}\right)^5 P_4(\cos \phi) &= \frac{35}{8} \left(\frac{r}{t}\right)^9 \sin^4 \theta - 5 \left(\frac{r}{t}\right)^7 \sin^2 \theta + \left(\frac{r}{t}\right)^5 \\ &= \sum_l \left(\frac{s}{r}\right)^{l-3} \left(\frac{1}{24} (1-x^2)^2 P_l'''_{l+1} - \frac{1}{3} (1-x^2) P_l''' + \frac{1}{3} P_l''\right) \\ &= \sum_l \left(\frac{s}{r}\right)^{l-4} \frac{1}{24} l(l-1)(l-2)(l-3) P_l. \end{aligned} \quad (\text{B5})$$

Thus we have from (B4) and (B5)

$$\frac{P_4(\cos \phi)}{t^3} = \sum_l \left(\frac{s^{l-4}}{r^{l-1}} \frac{7l(l-1)(l-2)(l-3)}{24(2l-1)} - \frac{s^{l-2}}{r^{l+1}} \frac{l(l-1)(l-2)(7l+11)}{24(2l+3)}\right) P_l(\cos \theta), \quad (\text{B6})$$

and

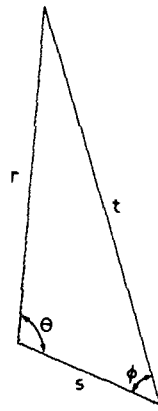


FIG. 3. General geometry considered in Appendix B for transforming the arguments of the Legendre polynomials from  $\cos \phi$  to  $\cos \theta$ .

$$\left(\frac{r}{t}\right)^9 = \frac{1}{105} \sum_l \left(\frac{s}{r}\right)^{l-4} P_l''''(\cos \theta). \quad (\text{B1e})$$

Next the law of sines gives

$$\sin \phi = (r/t) \sin \theta, \quad (\text{B2a})$$

and this along with  $t^2 = r^2 + s^2 - 2rs \cos \theta$  gives

$$\cos \phi = (r/t)(s/r - \cos \theta). \quad (\text{B2b})$$

For brevity, from now on we will write  $x = \cos \theta$  and drop the argument  $x$  on  $P_l$  and its derivatives; their presence will be assumed implicitly unless otherwise stated [as in  $P_l''(\cos \theta)$ ]. Also, we will freely use Legendre's equation (along with its two higher-order derivatives) and recurrence relations<sup>30</sup> for simplification. Similar relations for the associated Legendre functions,

$$P_l^m \equiv (1-x^2)^{m/2} d^m P_l(x)/dx^m, \quad (\text{B3})$$

will also be used whenever necessary.



$$\frac{P_4(\cos \phi)}{t^5} = \sum_T \frac{s^{l-4}}{r^{l+1}} \frac{1}{24} l(l-1)(l-2)(l-3) P_l(\cos \theta). \quad (\text{B7})$$

(ii)  $m = 1$  case

For the  $m = 1$  case we have

$$\begin{aligned} \left(\frac{r}{t}\right)^3 P_4^1(\cos \phi) &= \frac{5}{2} \sin \theta \left(\frac{s}{r} - \cos \theta\right) \left[4 \left(\frac{r}{t}\right)^5 - 7 \left(\frac{r}{t}\right)^7 \sin^2 \theta\right] \\ &= \frac{5}{2} (1-x^2)^{1/2} \sum_T \left(\frac{s}{r}\right)^{l-3} \left(\frac{4}{3} P_{l-2}'' - \frac{4}{3} x P_{l-1}'' - \frac{7}{15} (1-x^2) P_{l-1}''' + \frac{7}{15} x(1-x^2) P_{l+1}'''\right) \\ &= \sum_T \left(\frac{s}{r}\right)^{l-3} \frac{(l-2)}{6(2l+1)} [(7l+2)(l-3) P_{l-1}^1 - 7l(l-1) P_{l+1}^1] \\ &= \sum_T \left[ \left(\frac{s}{r}\right)^{l-4} \frac{(-7)(l-1)(l-2)(l-3)}{6(2l-1)} - \left(\frac{s}{r}\right)^{l-2} \frac{(-1)(l-1)(l-2)(7l+9)}{6(2l+3)} \right] P_l^1 \end{aligned} \quad (\text{B8})$$

and

$$\begin{aligned} \left(\frac{r}{t}\right)^5 P_4^1(\cos \phi) &= \frac{5}{2} \sin \theta \left(\frac{s}{r} - \cos \theta\right) \left[4 \left(\frac{r}{t}\right)^7 - 7 \left(\frac{r}{t}\right)^9 \sin^2 \theta\right] \\ &= (1-x^2)^{1/2} \sum_T \left(\frac{s}{r}\right)^{l-3} \left(\frac{2}{3} x P_{l-1}''' - \frac{2}{3} x P_l''' - \frac{1}{6} (1-x^2) P_l'''' + \frac{1}{6} x(1-x^2) P_{l+1}''''\right) \\ &= \sum_T \left(\frac{s}{r}\right)^{l-4} \left(-\frac{1}{6}\right) (l-1)(l-2)(l-3) P_l^1. \end{aligned} \quad (\text{B9})$$

Thus (B8) and (B9) give the desired relations

$$\frac{P_4^1(\cos \phi)}{t^3} = - \sum_T \left( \frac{s^{l-4}}{r^{l-1}} \frac{7(l-1)(l-2)(l-3)}{6(2l-1)} - \frac{s^{l-2}}{r^{l+1}} \frac{(l-1)(l-2)(7l+9)}{6(2l+3)} \right) P_l^1(\cos \theta) \quad (\text{B10})$$

and

$$\frac{P_4^1(\cos \phi)}{t^5} = - \sum_T \frac{s^{l-4}}{r^{l+1}} \frac{1}{6} (l-1)(l-2)(l-3) P_l^1(\cos \theta). \quad (\text{B11})$$

(iii)  $m = 2$  case

For the  $m = 2$  case we have

$$\begin{aligned} \left(\frac{r}{t}\right)^3 P_4^2(\cos \phi) &= \frac{15}{2} \sin^2 \theta \left[6 \left(\frac{r}{t}\right)^5 - 7 \left(\frac{r}{t}\right)^7 \sin^2 \theta\right] \\ &= \frac{15}{2} (1-x^2) \sum_T \left(\frac{s}{r}\right)^{l-3} \left(2 P_{l-1}'' - \frac{7}{15} (1-x^2) P_l'''\right) \\ &= \sum_T \left(\frac{s}{r}\right)^{l-3} \frac{1}{2(2l+1)} (7(l-1)(l-2) P_{l+1}^2 - (7l-4)(l-3) P_{l-1}^2) \\ &= \sum_T \left[ \left(\frac{s}{r}\right)^{l-4} \frac{7(l-2)(l-3)}{2(2l-1)} - \left(\frac{s}{r}\right)^{l-2} \frac{(l-2)(7l+3)}{2(2l+3)} \right] P_l^2 \end{aligned} \quad (\text{B12})$$

and

$$\begin{aligned} \left(\frac{r}{t}\right)^5 P_4^2(\cos \phi) &= \frac{15}{2} \sin^2 \theta \left[6 \left(\frac{r}{t}\right)^7 - 7 \left(\frac{r}{t}\right)^9 \sin^2 \theta\right] \\ &= (1-x^2) \sum_T \left(\frac{s}{r}\right)^{l-3} \left(3 P_l''' - \frac{1}{2} (1-x^2) P_{l+1}''''\right) \\ &= \sum_T \left(\frac{s}{r}\right)^{l-4} \frac{1}{2} (l-2)(l-3) P_l^2. \end{aligned} \quad (\text{B13})$$

Thus

$$\frac{P_4^2(\cos \phi)}{t^3} = \sum_T \left( \frac{s^{l-4}}{r^{l-1}} \frac{7(l-2)(l-3)}{2(2l-1)} - \frac{s^{l-2}}{r^{l+1}} \frac{(l-2)(7l+3)}{2(2l+3)} \right) P_l^2(\cos \theta), \quad (\text{B14})$$

and

$$\frac{P_4^2(\cos \phi)}{t^5} = \sum_T \frac{s^{l-4}}{r^{l+1}} \frac{1}{2} (l-2)(l-3) P_l^2(\cos \theta). \quad (\text{B15})$$

(iv)  $m = 3$  case

For the  $m = 3$  case we have

$$\begin{aligned} \left(\frac{r}{t}\right)^3 P_4^3(\cos \phi) &= 105 \sin^3 \theta \left(\frac{s}{r} - \cos \theta\right) \left(\frac{r}{t}\right)^7 \\ &= 7(1-x^2)^{3/2} \sum_T \left(\frac{s}{r}\right)^{l-3} (P_{l-1}''' - xP_l''') \\ &= \sum_T \left(\frac{s}{r}\right)^{l-3} \frac{7(l-2)}{2l+1} (P_{l-1}^3 - P_{l+1}^3) \\ &= - \sum_T \left[ \left(\frac{s}{r}\right)^{l-4} \frac{7(l-3)}{2l-1} - \left(\frac{s}{r}\right)^{l-2} \frac{7(l-1)}{2l+3} \right] P_l^3 \end{aligned} \quad (\text{B16})$$

and

$$\begin{aligned} \left(\frac{r}{t}\right)^5 P_4^3(\cos \phi) &= 105 \sin^3 \theta \left(\frac{s}{r} - \cos \theta\right) \left(\frac{r}{t}\right)^9 \\ &= (1-x^2)^{3/2} \sum_T \left(\frac{s}{r}\right)^{l-3} (P_l''' - xP_{l+1}''') \\ &= - \sum_T \left(\frac{s}{r}\right)^{l-4} (l-3) P_l^3. \end{aligned} \quad (\text{B17})$$

Thus

$$\begin{aligned} \frac{P_4^3(\cos \phi)}{t^3} &= - \sum_T \left( \frac{s^{l-4}}{r^{l-1}} \frac{7(l-3)}{2l-1} \right. \\ &\quad \left. - \frac{s^{l-2}}{r^{l+1}} \frac{7(l-1)}{2l+3} \right) P_l^3(\cos \theta) \end{aligned} \quad (\text{B18})$$

and

$$\frac{P_4^3(\cos \phi)}{t^5} = - \sum_T \frac{s^{l-4}}{r^{l+1}} (l-3) P_l^3(\cos \theta). \quad (\text{B19})$$

(v)  $m = 4$  case

For the  $m = 4$  case we have

$$\begin{aligned} \left(\frac{r}{t}\right)^3 P_4^4(\cos \phi) &= 105 \left(\frac{r}{t}\right)^7 \sin^4 \theta \\ &= 7(1-x^2)^2 \sum_T \left(\frac{s}{r}\right)^{l-3} P_l''' \\ &= \sum_T \left(\frac{s}{r}\right)^{l-3} \frac{7}{2l+1} (P_{l+1}^4 - P_{l-1}^4) \\ &= \sum_T \left[ \left(\frac{s}{r}\right)^{l-4} \frac{7}{(2l-1)} \right. \\ &\quad \left. - \left(\frac{s}{r}\right)^{l-2} \frac{7}{2l+3} \right] P_l^4 \end{aligned} \quad (\text{B20})$$

and

$$\begin{aligned} \left(\frac{r}{t}\right)^5 P_4^4(\cos \phi) &= 105 \left(\frac{r}{t}\right)^9 \sin^4 \theta \\ &= (1-x^2)^2 \sum_T \left(\frac{s}{r}\right)^{l-4} P_l''' \\ &= \sum_T \left(\frac{s}{r}\right)^{l-4} P_l^4. \end{aligned} \quad (\text{B21})$$

Thus

$$\frac{P_4^4(\cos \phi)}{t^3} = \sum_T \left( \frac{s^{l-4}}{r^{l-1}} \frac{7}{2l-1} - \frac{s^{l-2}}{r^{l+1}} \frac{7}{2l+3} \right) P_l^4(\cos \theta) \quad (\text{B22})$$

and

$$\frac{P_4^4(\cos \phi)}{t^5} = \sum_T \frac{s^{l-4}}{r^{l+1}} P_l^4(\cos \theta). \quad (\text{B23})$$

<sup>1</sup>G. K. Batchelor, *Annu. Rev. Fluid Mech.* **6**, 227 (1974).

<sup>2</sup>D. K. Hale, *J. Mater. Sci.* **11**, 2105 (1976).

<sup>3</sup>R. Landauer, in *Electrical Transport and Optical Properties of Inhomogeneous Media*, edited by J. C. Garland and D. B. Tanner (AIP, New York, 1978).

<sup>4</sup>Z. Hashin, *J. Appl. Mech.* **50**, 481 (1983).

<sup>5</sup>G. W. Milton, *Commun. Math. Phys.* **99**, 463 (1985).

<sup>6</sup>S. Torquato, *J. Appl. Phys.* **58**, 3790 (1985).

<sup>7</sup>S. Torquato, *Rev. Chem. Eng.* (in press). This review paper includes comparison of theoretical results to experimental data.

<sup>8</sup>S. Torquato and G. Stell, *J. Chem. Phys.* **77**, 2071 (1982).

<sup>9</sup>S. Torquato and G. Stell, *J. Chem. Phys.* **82**, 980 (1985).

<sup>10</sup>S. Torquato, *J. Stat. Phys.* **45**, 843 (1986).

<sup>11</sup>These self-consistent schemes are generally known as effective medium approximation with its many variants such as the coherent potential approximation, etc. For more details on them see Refs. 3 and 5 above. The name itself suggests that they mostly involve a mean-field approximation, though (in the context of a microscopic Hamiltonian) there exists a good self-consistent scheme to go beyond mean-field effects, namely the traveling cluster approximation (TCA). For more details regarding the usage of TCA in the microscopic context, see Refs. 12 and 13 below.

<sup>12</sup>R. Mills and P. Ratanavararaks, *Phys. Rev. B* **18**, 5291 (1978).

<sup>13</sup>Asok K. Sen, R. Mills, T. Kaplan, and L. J. Gray, *Phys. Rev. B* **30**, 5686 (1984).

<sup>14</sup>Z. Hashin and S. Shtrikman, *J. Mech. Phys. Solids* **11**, 127 (1963).

<sup>15</sup>J. J. McCoy, *Recent Advances in Engineering Science* (Gordon and Breach, New York, 1970), Vol. 5, pp. 235-254.

<sup>16</sup>G. W. Milton, *Phys. Rev. Lett.* **46**, 542 (1981).

<sup>17</sup>G. W. Milton and N. Phan-Thien, *Proc. R. Soc. London Ser. A* **380**, 305 (1982).

<sup>18</sup>S. Torquato, G. Stell, and J. Beasley, *Lett. Appl. Eng. Sci.* **23**, 385 (1985).

<sup>19</sup>S. Torquato, F. Lado, and P. Smith, *J. Chem. Phys.* **86**, 6388 (1987).

<sup>20</sup>F. Lado and S. Torquato, *Phys. Rev. B* **33**, 3370 (1986); S. Torquato and F. Lado, *ibid.* **33**, 6428 (1986).

<sup>21</sup>S. Torquato and J. D. Beasley, *Phys. Fluids* **30**, 633 (1987).

<sup>22</sup>Note that in Refs. 8 and 9,  $S_n$  denotes the  $n$ -point probability for the matrix. In the present work,  $S_n$  denotes the corresponding quantity for the particle phase.

<sup>23</sup>Note that here we have taken out the density factors explicitly instead of

keeping them implicitly in the order of the diagram as in Ref. 9.

<sup>24</sup>J. A. Barker and J. J. Monaghan, *J. Chem. Phys.* **36**, 2564 (1962). See also D. Henderson, *ibid.* **46**, 4306 (1967) and A. D. J. Haymet, S. A. Rice, and W. G. Madden, *ibid.* **74**, 3033 (1981) for similar applications.

<sup>25</sup>A. R. Edmonds, *Angular Momentum in Quantum Mechanics* (Princeton University Press, Princeton, New Jersey, 1957), Chap. 4.

<sup>26</sup>M. Abramowitz and I. A. Stegun, eds., *Handbook of Mathematical Func-*

*tions* (U. S. Government Printing Office, Washington, DC, 1964), Chap. 11.

<sup>27</sup>S. Torquato, *J. Chem. Phys.* **83**, 4776 (1985).

<sup>28</sup>Ref. 26, p. 440.

<sup>29</sup>Ref. 26, Chap. 22.

<sup>30</sup>See, for example, E. Jahnke and F. Emde, *Tables of Functions*, 4th ed. (Dover, New York, 1945), Chap. VII.